Dirac Equation in Very Special Relativity and the Gyromagnetic Factor

<u>Alessandro Santoni</u>^{1*}, Enrique Muñoz^{1†}, Benjamin Koch^{2**}

¹Pontificia Universidad Católica de Chile, Vicuña Mackenna 4860, Santiago, Chile ²Technische Universität Wien, Wiedner Hauptstrasse 8–10, A-1040 Vienna, Austria *asantoni@uc.cl, †munozt@fis.puc.cl, **benjamin.koch@tuwien.ac.at

Introducción

Very Special Relativity (VSR) [1] is a theory where flat spacetime symmetries are reduced to a subgroup of the Lorentz group plus the spacetime translations, which it is kept unchanged. While its kynematical consequences are identical to the ones of special relativity [1], new nonclassical consequences could arise. The original idea that motivated Cohen and Glashow to formulate VSR, for example, was a new mechanism for the emergence of Neutrino's masses [2].

In this work we consider the corrections arising from the SIM(2) invariant realization of Very Special Relativity to the energy spectrum of a C-invariant Dirac Fermion in a static and homogeneous magnetic field \vec{B} . First, we analyze the case of \vec{B} parallel to the spatial VSR preferred direction \vec{n} , finding that the expression for the energy spectrum stays the same, except for a mass shift $m \to m_f = \sqrt{m^2 + M^2}$. Then, we relax the parallelism condition, finding a new equation for the energy spectrum. We solve this equation perturbatively. With a Penning trap's experiment in mind, we derive the first order VSR corrections to the electron's g - 2factor. Finally, using the most accurate electron's g-factor measurements in Penning trap's experiments, we obtain an upper bound to the VSR electron mass parameter, and therefore also to the VSR electronic neutrino mass, of 1 eV.

Desarrollo

We start from the Modified VSR Dirac Equation in a fixed external magnetic field \vec{B}

$$\left(i\partial \!\!\!/ - eA - m + i\frac{M^2}{2}N\right)\psi(x) = 0, \qquad (1)$$

such that the VSR operator is $N^{\mu} = \frac{n^{\mu}}{n \cdot (\partial + ieA)}$, with n^{μ} representing the preferred VSR lightlike spacetime direction. From this we are able to find the perturbated energy spectrum in the general case, which, thinking of a trapped electron, can be related to the electron's gyromagnetic factor obtaining for its experimental-theoretical discrepancy

$$g_{exp} - g \sim -\mu \left[1 - \frac{11}{8} (2 - \frac{34}{11} \sin^2 \theta) \epsilon \right] \sin^2 \theta ,$$
 (2)

with $\mu = M/m_f$ and $\epsilon = eB/m_f^2$. Finally, considering the most up-to-date values of the theoretical and experimental value, we are able to give an upperbound to the VSR parameter of $\mu < 2.6 \times 10^{-12} \rightarrow M < 1 eV$.

Agradecimientos: A.S. acknowledges financial support from ANID Fellowship CONICYT-PFCHA/DoctoradoNacional/2020-21201387. E.M. acknowledges financial support from Fondecyt Grant No 1190361 and ANID PIA Anillo ACT192023.

Referencias

[1] A. G. Cohen and S. L. Glashow, Very special relativity, Physical review letters 97, 021601 (2006).

[2] A. G. Cohen and S. L. Glashow, A lorentz-violating ori- gin of neutrino mass?, arXiv preprint hep-ph/0605036 (2006).